

# **Dielectronic Recombination in O<sup>4+</sup> Near the Ionization Threshold**

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**ABSTRACT:** Relativistic perturbation theory calculations are carried out for dielectronic recombination cross sections involving levels in the  $O^{3+}$  1s<sup>2</sup>2p<sup>2</sup>3l (l = 0, 1, 2) configurations. One of the 8 levels in the 1s<sup>2</sup>2p<sup>2</sup>3s configuration is found to lie above the  $O^{4+}$  ionization limit, eight of the 21 levels in the 1s<sup>2</sup>2p<sup>2</sup>3p configuration are found to lie above the  $O^{4+}$  ionization limit, and all 28 levels in the 1s<sup>2</sup>2p<sup>2</sup>3d configuration are found to lie above the  $O^{4+}$  ionization limit. The largest cross sections are found at -2.0 eV for the 1s<sup>2</sup>2p<sup>2</sup>3s configuration, at 1.2 eV for the 1s<sup>2</sup>2p<sup>2</sup>3p configuration, and at 1.7 eV for the 1s<sup>2</sup>2p<sup>2</sup>3d configuration.

## 1. INTRODUCTION

Resonance states just below the ionization threshold may contribute to dielectronic recombination in astrophysical and laboratory plasmas[1, 2]. Proper inclusion of the below-threshold resonance states may make substantial changes in the thermally averaged rate coefficients.

Recently relativistic perturbation theory calculations were made for the dielectronic recombination cross sections involving the 10 levels in the  $F^{5+}$  1s<sup>2</sup>2p6p configuration[3]. Only one of the 10 levels was found to lie above the  $F^{6+}$  ionization limit. In this paper we look at dielectronic recombination in O<sup>4+</sup> 1s<sup>2</sup>2s<sup>2</sup> leading to O<sup>3+</sup> 1s<sup>2</sup>2p<sup>2</sup>3l (l = 0, 1, 2) resonance states that lie just below and above the ionization threshold.

The rest of the paper is structured as follows: in section 2 we review theory, in section 3 we present results, and in section we give a brief summary. Unless otherwise stated, all quantities are given in atomic units.

## 2. THEORY

The dielectronic recombination cross section for an N electron ground level with statistical weight  $g_i$  combining into an (N+1) electron doubly excited level with statistical weight  $g_i$  is given by[4]:

$$\sigma_{i \to j} = \frac{\pi^2}{E_c \Delta E_c} \frac{g_j}{2g_i} A_a(j \to i) B_j \tag{1}$$

where  $E_c$  is the energy of the continuum electron and  $\Delta E_c$  is bin width. The branching ratio for radiative stabilization is given by:

$$B_j = \frac{\sum_n A_r(j \to n)}{\sum_k A_a(j \to k) + \sum_n A_r(j \to n)}$$
(2)

where the radiative A<sub>r</sub> and autoionization A<sub>a</sub> rates are evaluated using lowest order perturbation theory.

For radiative and autoionization rates, the energies and bound state wavefunctions are calculated using a multi-

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configuration Dirac-Fock atomic structure code[5]. The continuum state wavefunctions are calculated using a a single-channel radial Dirac equation, where the Dirac local exchange distorting potential is constructed from Dirac-Fock bound radial orbitals.

# 3. RESULTS

Fully relativistic calculations were carried out for the O<sup>4+</sup>  $1s^22s^2$  subconfiguration leading to an energy of -68.3032 au. Fully relativistic calculations were also carried out for the O<sup>4+</sup>  $(1s^22s^2+1s^22p^2+1s^22p^2)$  subconfigurations leading to an energy of -68.3970 au. Energies for the 57 levels associated with the O<sup>3+</sup>  $1s^22p^2$  (1 = 0, 1, 2) configurations are based on differences between low-order relativistic calculations for the resonance subconfigurations and fully-relativistic calculations for the O<sup>4+</sup>  $(1s^22s^2+1s^22p^2)$  subconfigurations.

Low-order relativistic calculations were carried out for the 8 levels of the O<sup>3+</sup> even-parity subconfigurations:  $1s^{2}2p^{2}3s$  (1 level),  $1s22p^{2}2p^{2}3s$  (4 levels), and  $1s^{2}2p^{2}3s$  (3 levels). For added correlation we also included 9 levels from the  $1s^{2}2s^{2}p^{2}$  and  $1s^{2}2s^{2}3s$  configurations. Autoionization rates were determined for the decay of the 8 levels of the  $1s^{2}2p^{2}3s$  configuration to the 30 levels of the  $1s^{2}2s^{2}pkp$  and  $1s^{2}2s^{2}pkf$  configurations. Radiative rates were determined for the decay of the 8 levels of the  $1s^{2}2p^{2}3s$  configuration to the 30 levels of the  $1s^{2}2p^{2}3s$  configurations. Radiative rates were determined for the decay of the 8 levels of the  $1s^{2}2p^{2}3s$  configuration to the 5 levels of the  $1s^{2}2p^{3}$  configuration. Energies, autoionization rates, and radiative rates for the 8 levels are presented in Table 1. Dielectronic recombination cross sections for the 8 levels are presented in Figure 1 using  $\Delta E_{c} = 0.04$  eV in Eq.(1). Convoluted dielectronic recombination cross sections for the 8 levels are presented in Figure 2 using a convolution energy of 0.4 eV. We note that the largest convoluted cross section is around -2.0 eV due to the levels  $1s^{2}2p^{2}(0)3s$  1/2 and  $1s^{2}2p^{2}(2)3s$  3/2.

Low-order relativistic calculations were carried out for the 28 levels of the O<sup>3+</sup> even-parity subconfigurations:  $1s^{2}2p^{2}3d(1 \text{ level})$ ,  $1s^{2}2p^{2}2p3d(7 \text{ levels})$ ,  $1s^{2}2p^{2}3d(5 \text{ levels})$ ,  $1s^{2}2p^{2}3d(1 \text{ level})$ ,  $1s^{2}2p^{2}2p^{2}d(8 \text{ levels})$ , and  $1s^{2}2p^{2}3d(6 \text{ levels})$ . For added correlation we also included 10 levels from the  $1s^{2}2s^{2}p^{2}$  and  $1s^{2}2s^{2}3d$  configurations. Autoionization rates were determined for the decay of the 28 levels of the  $1s^{2}2p^{2}3d$  configuration to the 53 levels of the  $1s^{2}2s^{2}pkp$ ,  $1s^{2}2s^{2}pkf$ , and  $1s^{2}2s^{2}pkh$  configurations. Radiative rates were determined for the decay of the 28 levels of the  $1s^{2}2p^{2}3d$  configuration to the 5 levels of the  $1s^{2}2p^{3}$  configuration. Energies, autoionization rates, and radiative rates for the 28 levels are presented in Table 3. Dielectronic recombination cross sections for the 28 levels are presented in Figure 5 using  $\Delta E_{c} = 0.04$  eV in Eq.(1). Convoluted dielectronic recombination cross sections for the 28 levels are presented in Figure 6 using a convolution energy of 0.4 eV. We note that the largest convoluted cross section is around 1.7 eV due to the levels  $1s^{2}2p^{2}(2)3d(3/2)$ ,  $1s^{2}2p^{2}(2)3d(5/2)$ , and  $1s^{2}2p^{2}(2)3d(7/2)$ .

# 4. SUMMARY

Relativistic perturbation theory has been applied to calculate the dielectronic recombination in  $O^{4+}$  above and below the ionization threshold involving levels in the  $O^{3+}$  1s<sup>2</sup>2p<sup>2</sup>3s (l = 0, 1, 2) configurations. In the future we plan to map out the dielectronic recombination above and below the ionization threshold for the  $O^{3+}$  1s<sup>2</sup>2s2p4l (l = 0, 1, 2, 3) and 1s<sup>2</sup>2s2p5l (l = 0, 1, 2, 3, 4) configurations.

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Sub-configuration	J value	Energy	Auger Rate	Radiative Rate
$1s^2 2\bar{p}^2(0)3s$	1/2	-5.50  eV	$3.17 \times 10^{14} \text{ Hz}$	$2.66 \times 10^9 \text{ Hz}$
$1s^22\bar{p}2p(1)3s$	1/2	-5.48  eV	$3.22 \times 10^{14} \text{ Hz}$	$2.62 \times 10^9 \text{ Hz}$
$1s^22\bar{p}2p(1)3s$	3/2	-5.44  eV	$1.37 \times 10^{14} \text{ Hz}$	$2.56 \times 10^9 \text{ Hz}$
$1s^22\bar{p}2p(2)3s$	3/2	-4.18  eV	$5.94 \times 10^{15} \text{ Hz}$	$5.27 \times 10^9 \text{ Hz}$
$1s^22\bar{p}2p(2)3s$	5/2	-4.14  eV	$7.39\times10^{15}~{\rm Hz}$	$5.27 \times 10^9 \text{ Hz}$
$1s^22p^2(0)3s$	1/2	-1.93  eV	$9.63 \times 10^{14} \text{ Hz}$	$4.11 \times 10^9 \text{ Hz}$
$1s^22p^2(2)3s$	3/2	-1.93  eV	$1.76 \times 10^{15} \text{ Hz}$	$4.19 \times 10^9 \text{ Hz}$
$1s^22p^2(2)3s$	5/2	$5.17 \ \mathrm{eV}$	$1.00 \times 10^{15} \text{ Hz}$	$5.60 \times 10^9 \text{ Hz}$

Table 1. O<sup>3+</sup> 1s<sup>2</sup>2p<sup>2</sup>3s Auger and Radiative Rates

Table 2. O<sup>3+</sup> 1s<sup>2</sup>2p<sup>2</sup>3p Auger and Radiative Rates

Sub-configuration	J value	Energy	Auger Rate	Radiative Rate
$1s^2 2\bar{p}^2(0) 3\bar{p}$	1/2	-3.10  eV	$2.48 \times 10^{11} \text{ Hz}$	$4.93 \times 10^9 \text{ Hz}$
$1s^22\bar{p}2p(1)3\bar{p}$	1/2	-2.58  eV	$1.45\times10^{13}~{\rm Hz}$	$1.18 \times 10^{10} \text{ Hz}$
$1s^22\bar{p}2p(1)3\bar{p}$	3/2	-2.57  eV	$1.47\times10^{13}~{\rm Hz}$	$1.18\times10^{10}~{\rm Hz}$
$1s^22\bar{p}2p(2)3\bar{p}$	3/2	-2.54  eV	$1.49 \times 10^{13} \text{ Hz}$	$1.18 \times 10^{10} \text{ Hz}$
$1s^22\bar{p}2p(2)3\bar{p}$	5/2	-2.50  eV	$1.52\times10^{13}~{\rm Hz}$	$1.17\times10^{10}~{\rm Hz}$
$1s^22p^2(0)3\bar{p}$	1/2	-2.18  eV	$1.16\times10^{14}~{\rm Hz}$	$1.20 \times 10^{10} \text{ Hz}$
$1s^22p^2(2)3\bar{p}$	3/2	-2.17  eV	$1.15\times10^{14}~{\rm Hz}$	$1.20\times10^{10}~{\rm Hz}$
$1s^22p^2(2)3\bar{p}$	5/2	-2.15  eV	$1.14\times10^{14}~{\rm Hz}$	$1.20 \times 10^{10} \text{ Hz}$
$1s^22\bar{p}^2(0)3p$	3/2	-1.60  eV	$2.09\times10^{13}~{\rm Hz}$	$5.87 \times 10^9 \text{ Hz}$
$1s^22\bar{p}2p(1)3p$	1/2	-1.56  eV	$2.00\times10^{13}~{\rm Hz}$	$5.89 \times 10^9 \text{ Hz}$
$1s^22\bar{p}2p(1)3p$	3/2	-1.37  eV	$1.57\times10^{11}~{\rm Hz}$	$1.26 \times 10^{10} \text{ Hz}$
$1s^22\bar{p}2p(1)3p$	5/2	-0.91  eV	$3.84 \times 10^{14} \text{ Hz}$	$5.93 \times 10^9 \text{ Hz}$
$1s^22\bar{p}2p(2)3p$	1/2	-0.90  eV	$3.81 \times 10^{14} \text{ Hz}$	$6.03 \times 10^9 \text{ Hz}$
$1s^22\bar{p}2p(2)3p$	3/2	1.16  eV	$1.86\times10^{13}~{\rm Hz}$	$9.37 \times 10^9 \text{ Hz}$
$1s^2 2\bar{p} 2p(2) 3p$	5/2	1.18  eV	$2.03 \times 10^{13} \text{ Hz}$	$9.31 \times 10^9 \text{ Hz}$
$1s^22\bar{p}2p(2)3p$	7/2	1.67  eV	$1.06 \times 10^{13} \text{ Hz}$	$9.24 \times 10^9 \text{ Hz}$
$1s^22p^2(0)3p$	3/2	1.68  eV	$1.07\times10^{13}~{\rm Hz}$	$9.26 \times 10^9 \text{ Hz}$
$1s^22p^2(2)3p$	1/2	2.36  eV	$7.02\times10^{14}~{\rm Hz}$	$8.64\times10^9~{\rm Hz}$
$1s^22p^2(2)3p$	3/2	2.39  eV	$7.20\times10^{14}~{\rm Hz}$	$8.65 \times 10^9 \text{ Hz}$
$1s^22p^2(2)3p$	5/2	8.88  eV	$1.58\times10^{14}~{\rm Hz}$	$1.01\times10^{10}~{\rm Hz}$
$1s^22p^2(2)3p$	7/2	8.89 eV	$1.57\times10^{14}~\mathrm{Hz}$	$1.01\times10^{10}~{\rm Hz}$

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<b>Table 3. O<sup>3+</sup></b> (	$1s^22p^23d$	Auger and	<b>Radiative Rates</b>
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Sub-configuration	J value	Energy	Auger Rate	Radiative Rate
$1s^2 2\bar{p}^2(0) 3\bar{d}$	3/2	0.64  eV	$2.17 \times 10^{16} \text{ Hz}$	$7.11 \times 10^5 \text{ Hz}$
$1s^2 2\bar{p}2p(1)3d$	1/2	0.65 eV	$2.09 \times 10^{16} \text{ Hz}$	$4.69 \times 10^5 \text{ Hz}$
$1s^2 2\bar{p}2p(1)3\bar{d}$	3/2	0.68 eV	$2.07 \times 10^{16} \text{ Hz}$	$2.23 \times 10^5 \text{ Hz}$
$1s^2 2\bar{p}2p(1)3\bar{d}$	5/2	0.70  eV	$2.08 \times 10^{16} \ \mathrm{Hz}$	0.00 Hz
$1s^2 2\bar{p} 2p(2) 3\bar{d}$	1/2	1.16 eV	$1.62 \times 10^{16} \text{ Hz}$	$6.92 \times 10^9 \text{ Hz}$
$1s^2 2\bar{p}2p(2)3d$	3/2	1.18  eV	$1.62 \times 10^{16} \text{ Hz}$	$4.18 \times 10^9 \text{ Hz}$
$1s^2 2\bar{p} 2p(2) 3\bar{d}$	5/2	1.22  eV	$1.74\times10^{16}~{\rm Hz}$	$6.77 \times 10^8 \text{ Hz}$
$1s^2 2\bar{p} 2p(2) 3\bar{d}$	7/2	1.22  eV	$1.73 \times 10^{16} \ \mathrm{Hz}$	$5.10 \times 10^7 \text{ Hz}$
$1s^22p^2(0)3d$	3/2	1.22  eV	$1.90 \times 10^{16} \text{ Hz}$	$3.41 \times 10^9 \text{ Hz}$
$1s^22p^2(2)3d$	1/2	1.24  eV	$1.89\times10^{16}~\mathrm{Hz}$	$3.53 \times 10^{6} \text{ Hz}$
$1s^2 2p^2(2) 3\bar{d}$	3/2	1.72  eV	$1.59 \times 10^{16} \ \mathrm{Hz}$	$4.35\times10^{10}~\mathrm{Hz}$
$1s^2 2p^2(2) 3\bar{d}$	5/2	1.74  eV	$1.58 \times 10^{16} \ \mathrm{Hz}$	$4.39\times10^{10}~{\rm Hz}$
$1s^22p^2(2)3d$	7/2	1.75  eV	$1.75\times10^{16}~{\rm Hz}$	$4.42\times10^{10}~\mathrm{Hz}$
$1s^2 2\bar{p}^2(0)3d$	5/2	1.83  eV	$1.56\times10^{16}~{\rm Hz}$	$7.99\times10^9~{\rm Hz}$
$1s^2 2\bar{p} 2p(1) 3d$	3/2	1.88  eV	$1.71\times10^{16}~{\rm Hz}$	$8.16 \times 10^9 \text{ Hz}$
$1s^2 2\bar{p}2p(1)3d$	5/2	3.45 eV	$1.06\times10^{16}~{\rm Hz}$	$2.85 \times 10^{10} \text{ Hz}$
$1s^2 2\bar{p}2p(1)3d$	7/2	3.47  eV	$9.96 \times 10^{15} \text{ Hz}$	$2.90\times10^{10}~\mathrm{Hz}$
$1s^2 2\bar{p} 2p(2) 3d$	1/2	4.51 eV	$9.30\times10^{15}~\mathrm{Hz}$	$9.52 \times 10^{6} \text{ Hz}$
$1s^2 2\bar{p} 2p(2) 3d$	3/2	4.52  eV	$9.32 \times 10^{15} \text{ Hz}$	0.00 Hz
$1s^2 2\bar{p}2p(2)3d$	5/2	4.92 eV	$9.22 \times 10^{15} \text{ Hz}$	$2.56\times10^{10}~{\rm Hz}$
$1s^2 2\bar{p}2p(2)3d$	7/2	4.93 eV	$8.25\times10^{15}~\mathrm{Hz}$	$2.54\times10^{10}~\mathrm{Hz}$
$1s^2 2\bar{p} 2p(2) 3d$	9/2	4.96 eV	$9.91 \times 10^{15} \text{ Hz}$	$4.74 \times 10^{10} \text{ Hz}$
$1s^22p^2(0)3d$	5/2	4.97  eV	$1.01 \times 10^{16} \text{ Hz}$	$4.76 \times 10^{10} \ \mathrm{Hz}$
$1s^22p^2(2)3d$	1/2	5.94  eV	$9.00 \times 10^{15} \text{ Hz}$	$1.72\times10^{10}~\mathrm{Hz}$
$1s^22p^2(2)3d$	3/2	6.02  eV	$9.70 \times 10^{15} \text{ Hz}$	$3.40 \times 10^{10} \text{ Hz}$
$1s^22p^2(2)3d$	5/2	6.02  eV	$9.71 \times 10^{15} \ \mathrm{Hz}$	$3.41\times10^{10}~\mathrm{Hz}$
$1s^22p^2(2)3d$	7/2	12.07  eV	$4.45 \times 10^{15} \ \mathrm{Hz}$	$2.13\times10^{10}~\mathrm{Hz}$
$1s^22p^2(2)3d$	9/2	12.08  eV	$5.33\times10^{15}~\mathrm{Hz}$	$2.20\times10^{10}~\mathrm{Hz}$



Figure 1. Dielectronic recombination cross sections for  $O^{4+}$  1s<sup>2</sup>2s<sup>2</sup> via  $O^{3+}$  1s<sup>2</sup>2p<sup>2</sup>3s (1.0 Mb = 1.0 × 10<sup>-18</sup> cm<sup>2</sup>).



Figure 2. Convoluted dielectronic recombination cross sections for  $O^{4+}$  1s<sup>2</sup>2s<sup>2</sup> via  $O^{3+}$  1s<sup>2</sup>2p<sup>2</sup>3s (1.0 Mb = 1.0 × 10<sup>-18</sup> cm<sup>2</sup>).



Figure 3. Dielectronic recombination cross sections for  $O^{4+}$  1s<sup>2</sup>2s<sup>2</sup> via  $O^{3+}$  1s<sup>2</sup>2p<sup>2</sup>3p (1.0 Mb = 1.0 × 10<sup>-18</sup> cm<sup>2</sup>).



Figure 4. Convoluted dielectronic recombination cross sections for  $O^{4+}$  1s<sup>2</sup>2s<sup>2</sup> via  $O^{3+}$  1s<sub>2</sub>2p<sup>2</sup>3p (1.0 Mb = 1.0 × 10<sup>-18</sup> cm<sup>2</sup>).



Figure 5. Dielectronic recombination cross sections for  $O^{4+}$  1s<sup>2</sup>2s<sup>2</sup> via  $O^{3+}$  1s<sup>2</sup>2p<sup>2</sup>3d (1.0 Mb = 1.0 × 10<sup>-18</sup> cm<sup>2</sup>).



Figure 6. Convoluted dielectronic recombination cross sections for  $O^{4+}$  1s<sup>2</sup>2s<sup>2</sup> via  $O^{3+}$  1s<sup>2</sup>2p<sup>2</sup>3d (1.0 Mb = 1.0 × 10<sup>-18</sup> cm<sup>2</sup>).