

Laboratory Micrometeoroid/Dust Ablation Studies

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Introduction

- Every day, billions of micrometeoroids ablate in Earth's upper atmosphere.
- The ablated materials affect a variety of phenomena:
 - Formation of layers of metal atoms and ions in the atmosphere
 - Nucleation of noctilucent clouds
 - Effects on stratospheric aerosols and O₃ chemistry
 - Informs us about the dust environment of the solar system

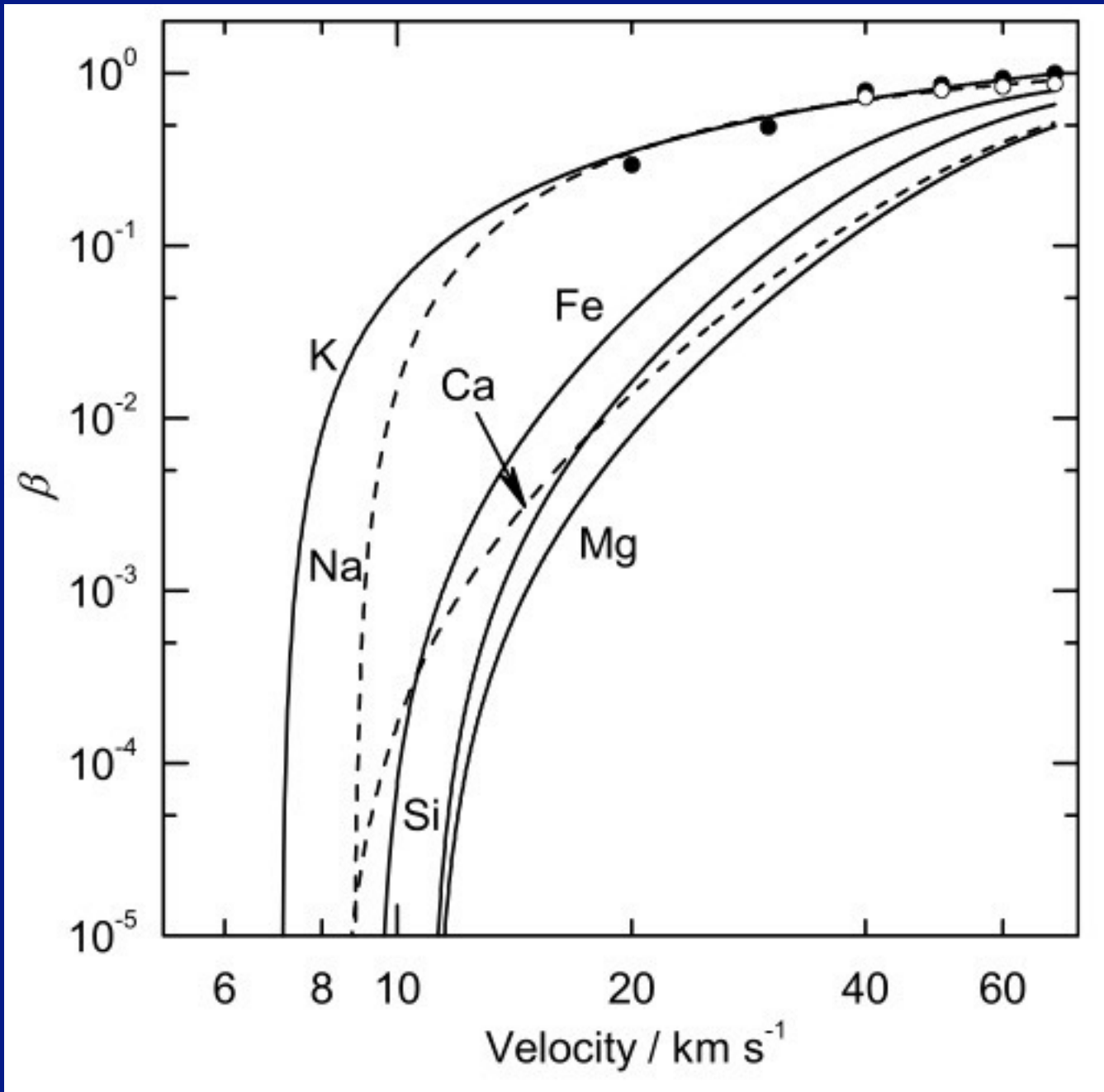
Introduction

- First laboratory measurements of the complete ablation process.
- Spatial and temporal resolution
- Extended measurements to relevant velocities.
- First measurements address ionization efficiency (β)
- Long term goal is detailed validation of ablation models.

Meteors Ionize the Air

- The meteor collides with gas molecules and begins to heat.
 - Sputtering
- The meteor melts and the most volatile elements begin to evaporate.
 - Evaporated atoms collide with gas molecules and can form ion/electron pairs.
 - Amount of ionization given by β .
 - As meteor continues to heat, less volatile elements evaporate, etc.
- This process is modeled in meteor models, such as the Chemical Ablation Model (CABMOD, Vondrak et al. 2008).

Ionization (β) is Uncertain Quantity in Models

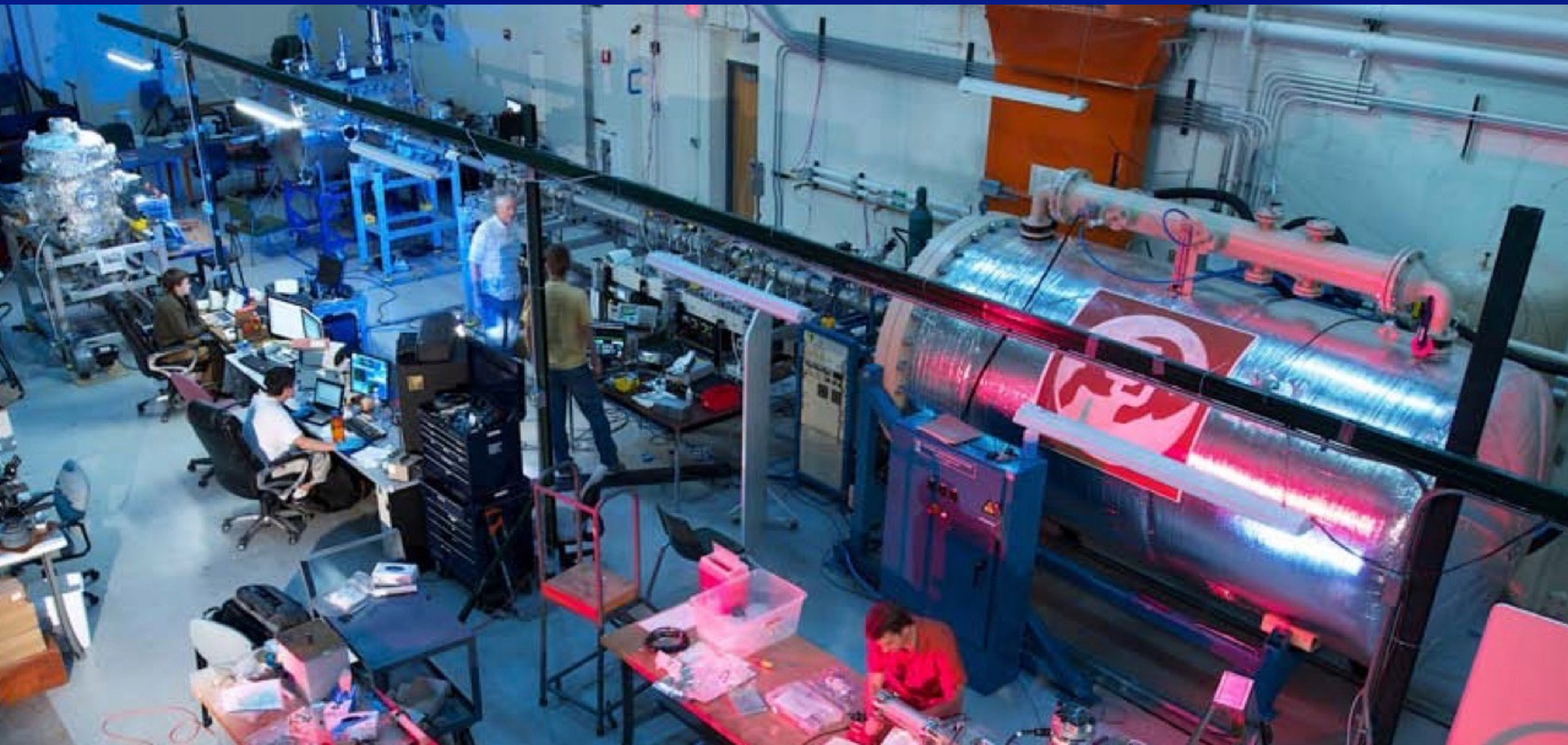


[Vondrak et al., 2008]

- Jones (1997) model of ionization coefficient is used in CABMOD.
- Very few measurements at low velocity.
- β is crucial for interpreting radar measurements of meteors.

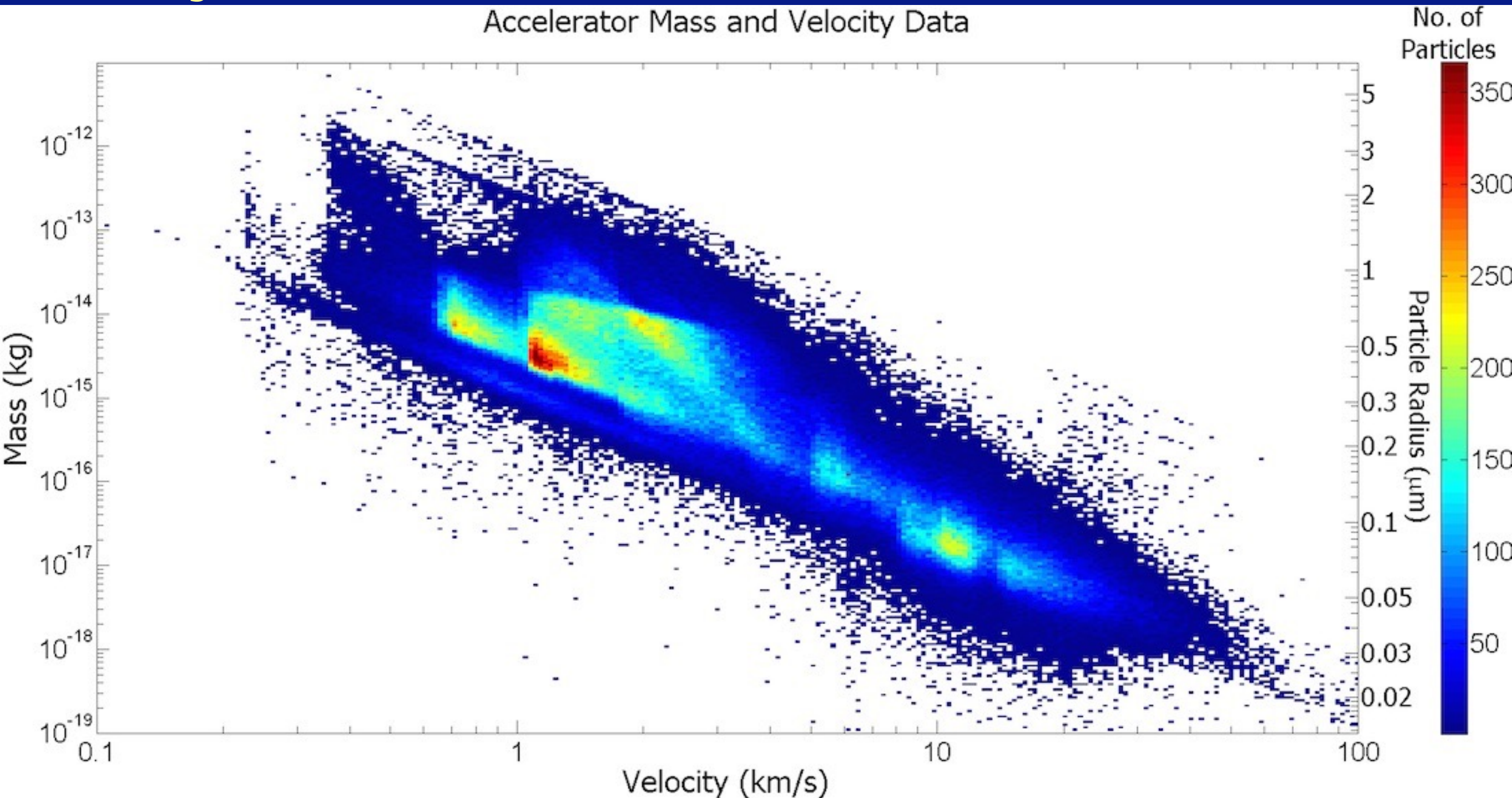
Experiment: Measure β

- Using a 3MV hypervelocity dust accelerator (Shu, 2012), we can simulate ablating micrometeoroids in the lab.
 - This experiment used micron and sub-micron sized iron particles.

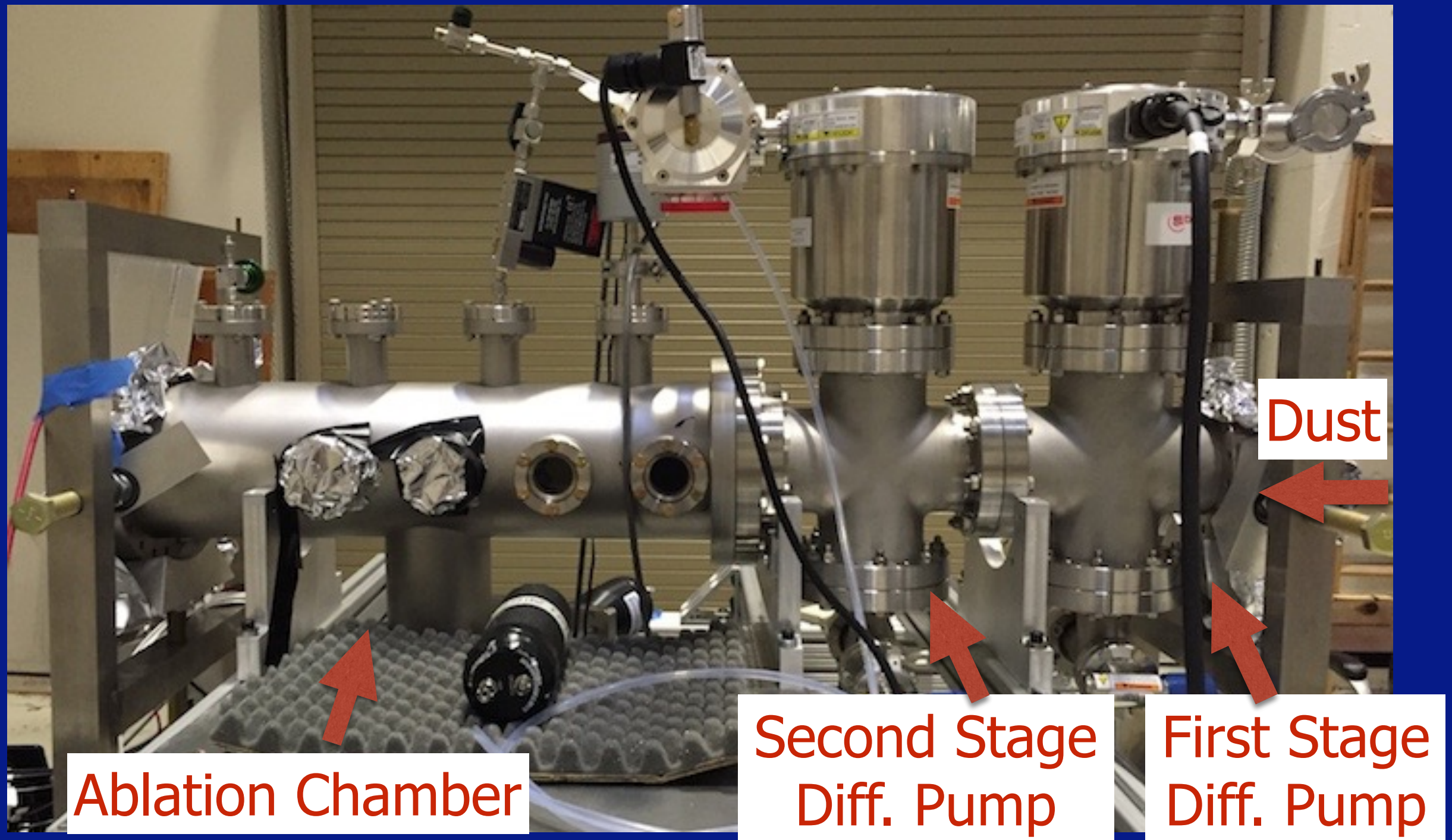


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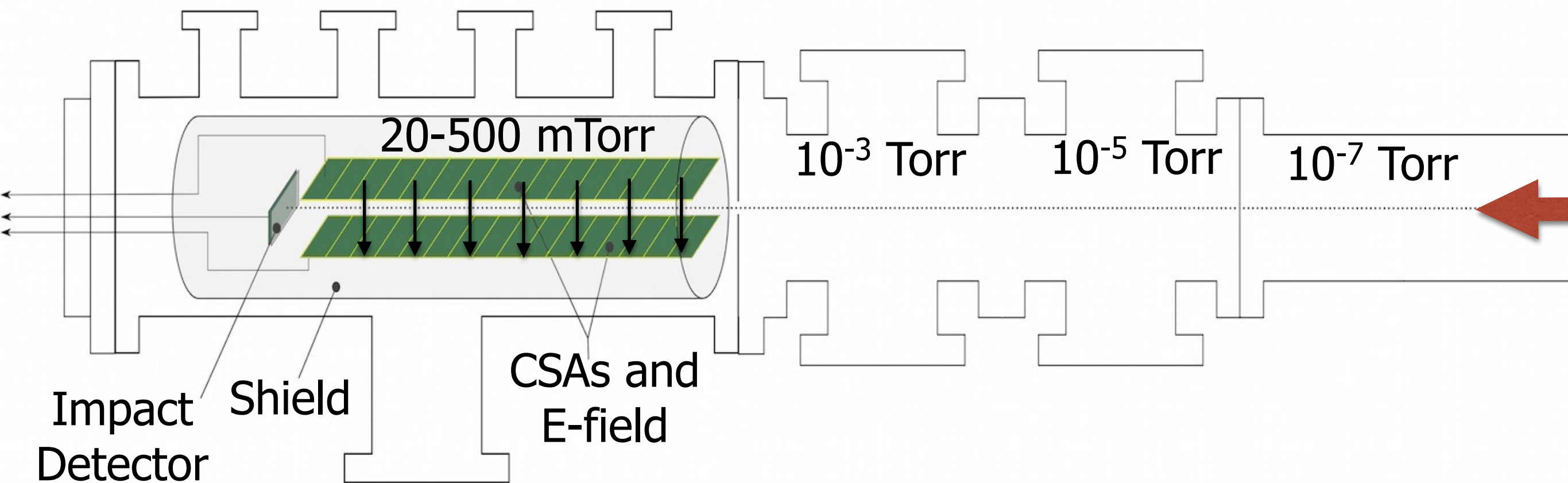


Experimental Apparatus



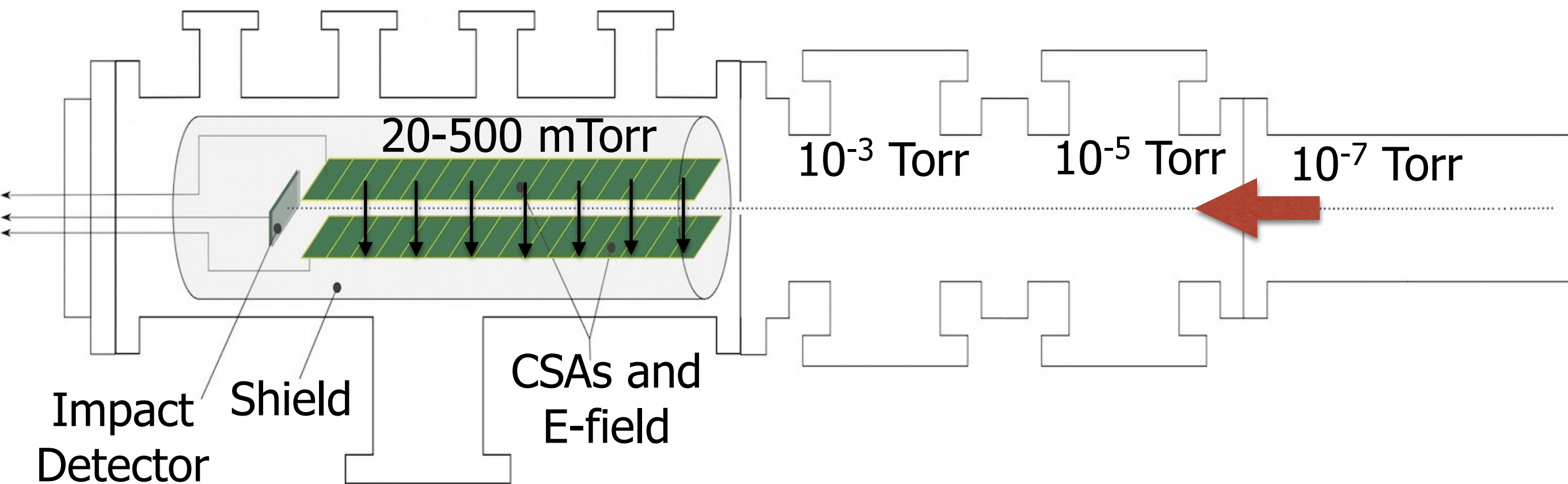
Experimental Apparatus

- The dust enters the experimental apparatus, shown below.



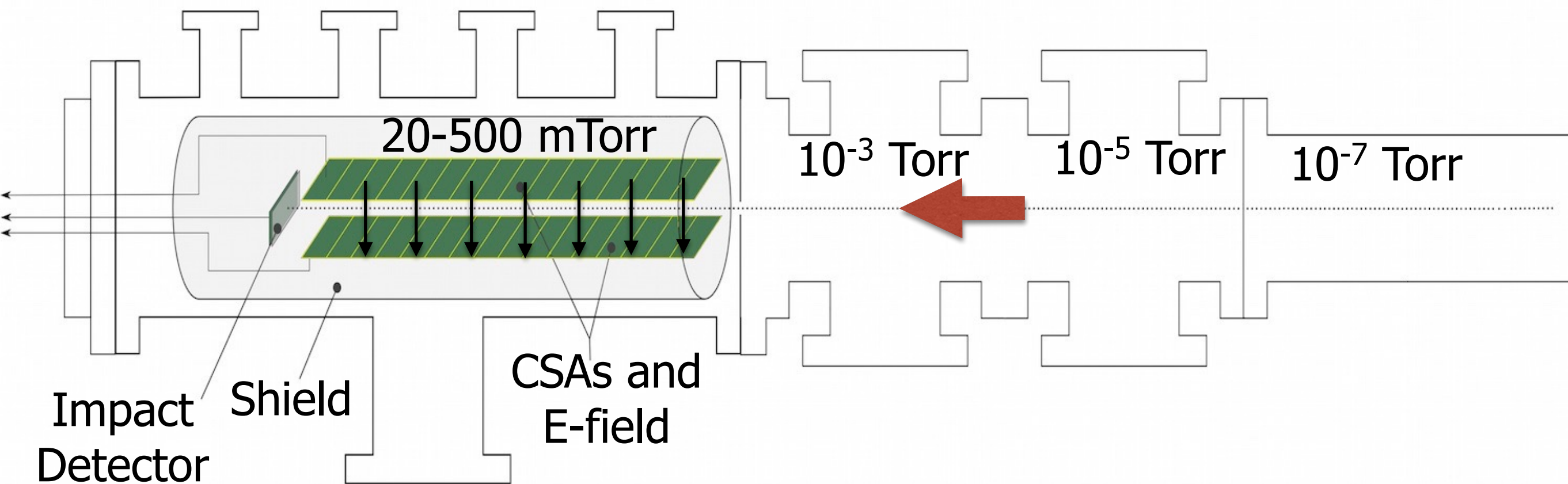
Experimental Apparatus

- The dust moves through two stages of differential pumping.



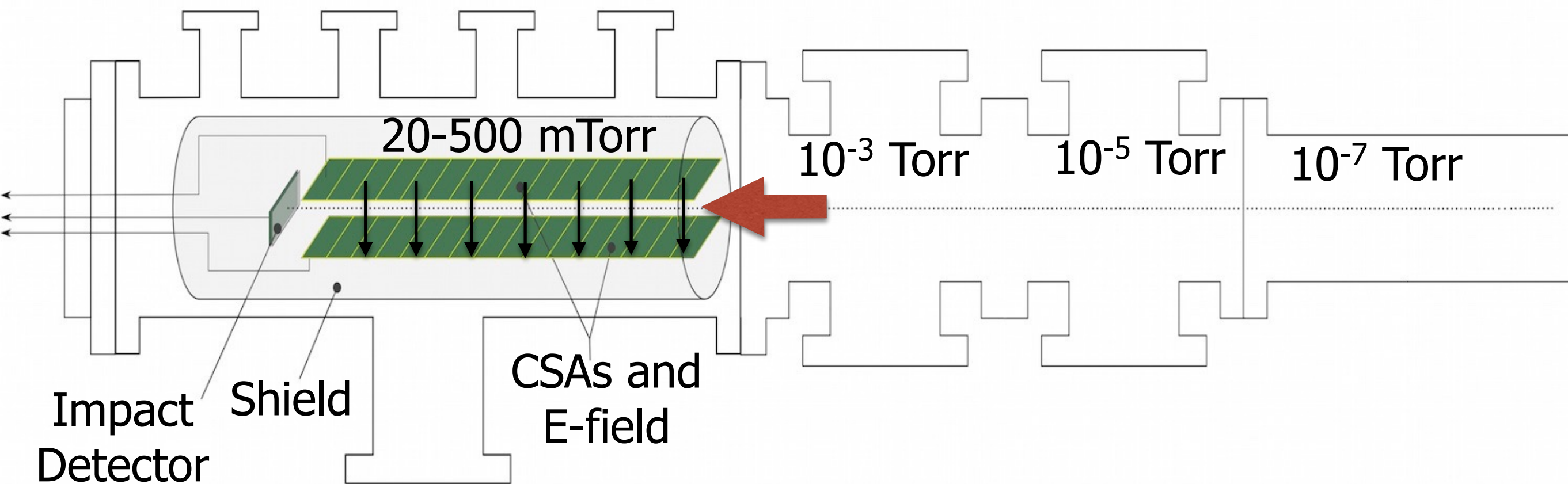
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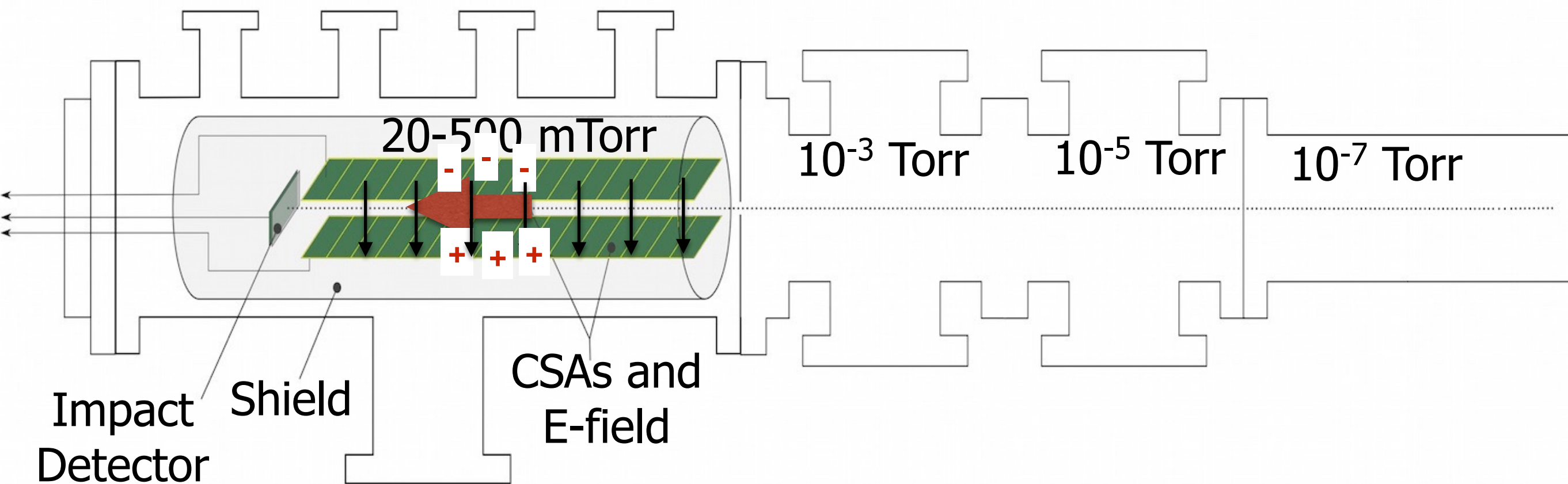
Experimental Apparatus

- The dust enters the ablation chamber with biased charge sensitive amplifiers (CSAs) above and below it.
 - The ablation gas used was N_2 .



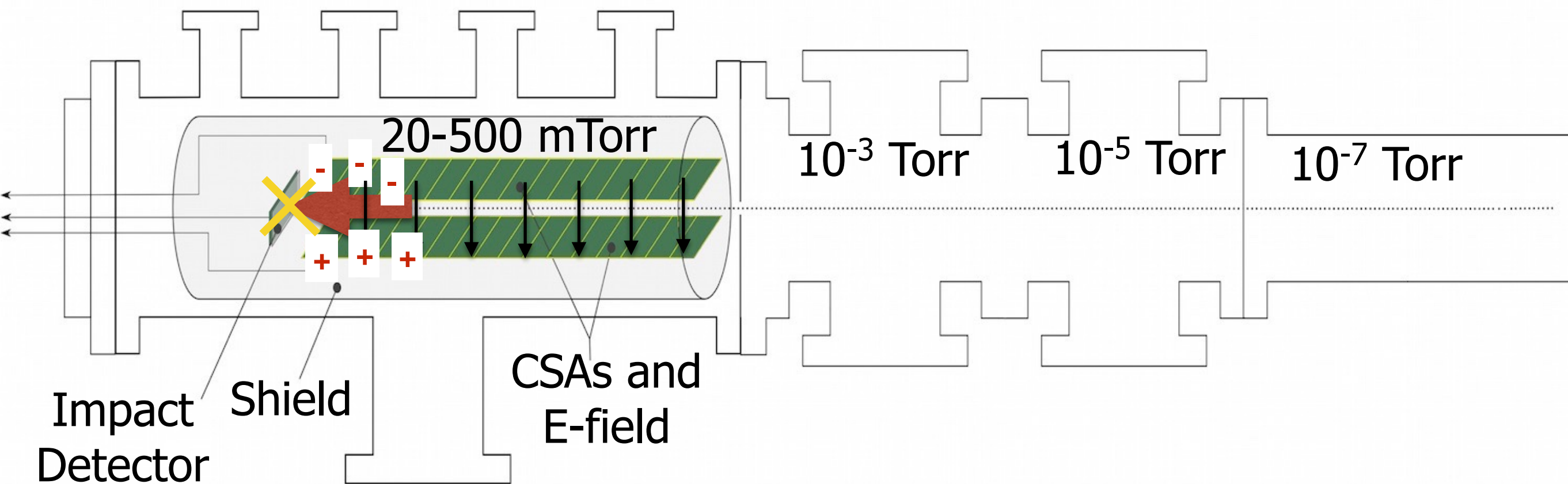
Experimental Apparatus

- Ions and electrons are collected separately on the top and bottom CSAs.

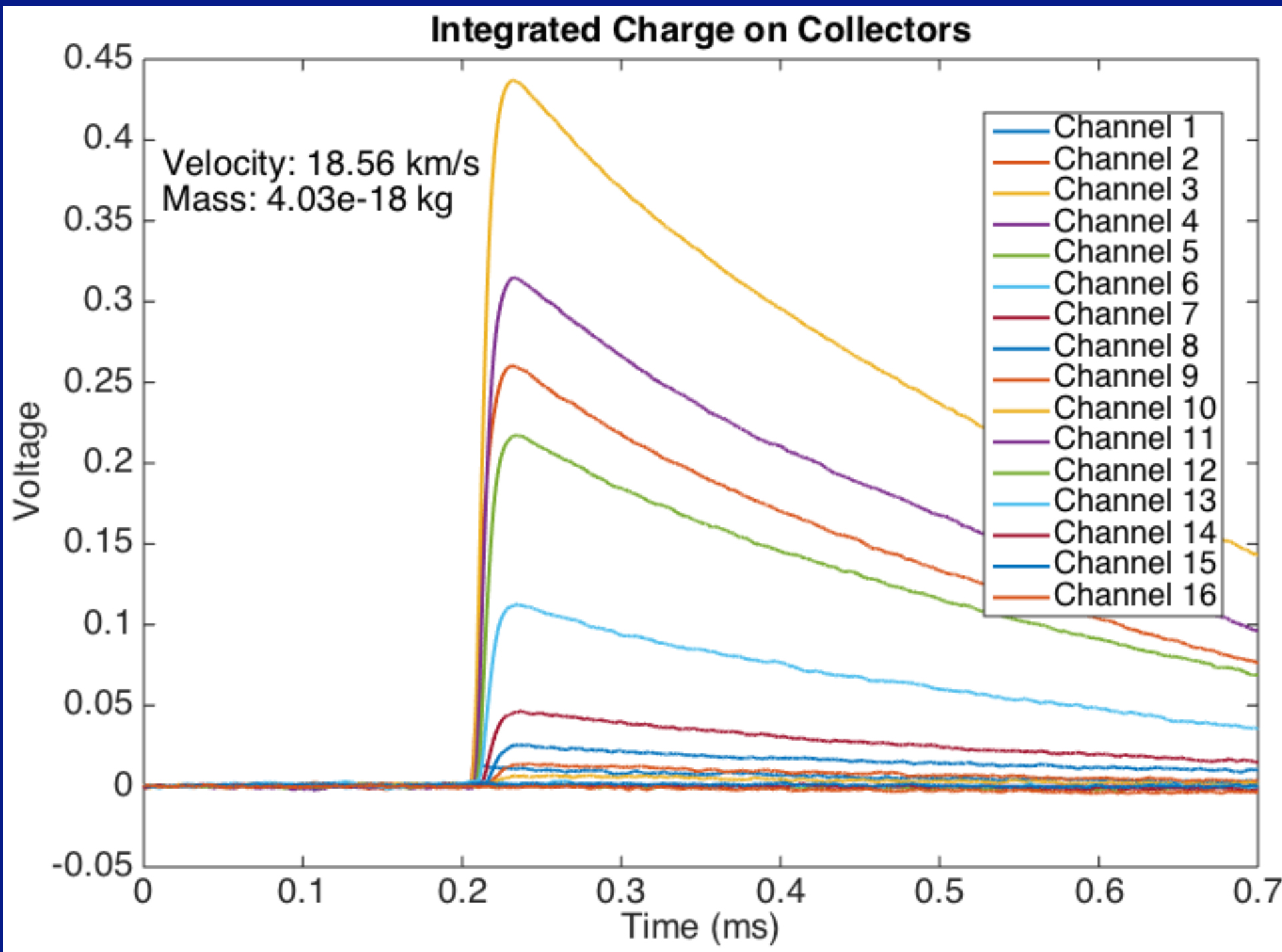


Experimental Apparatus

- If the particle survives, whatever is left of the particle strikes the impact detector.
 - We looked for particles with no impact signal.

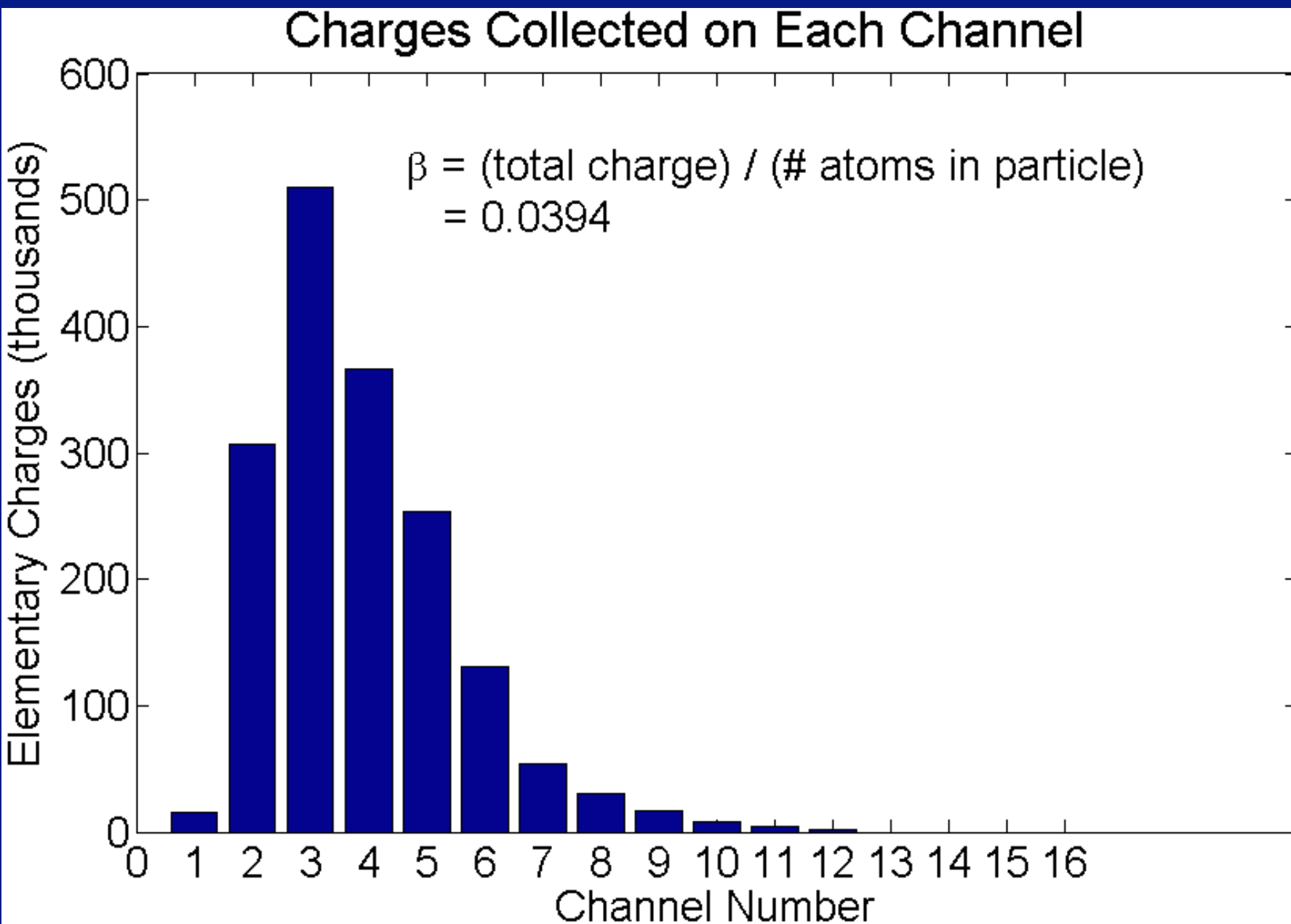


Signals Spatially/Temporally Separated



- Each CSA channel collects the charge which is produced near its collector.
- This results in temporally and spatially separated charge signals.

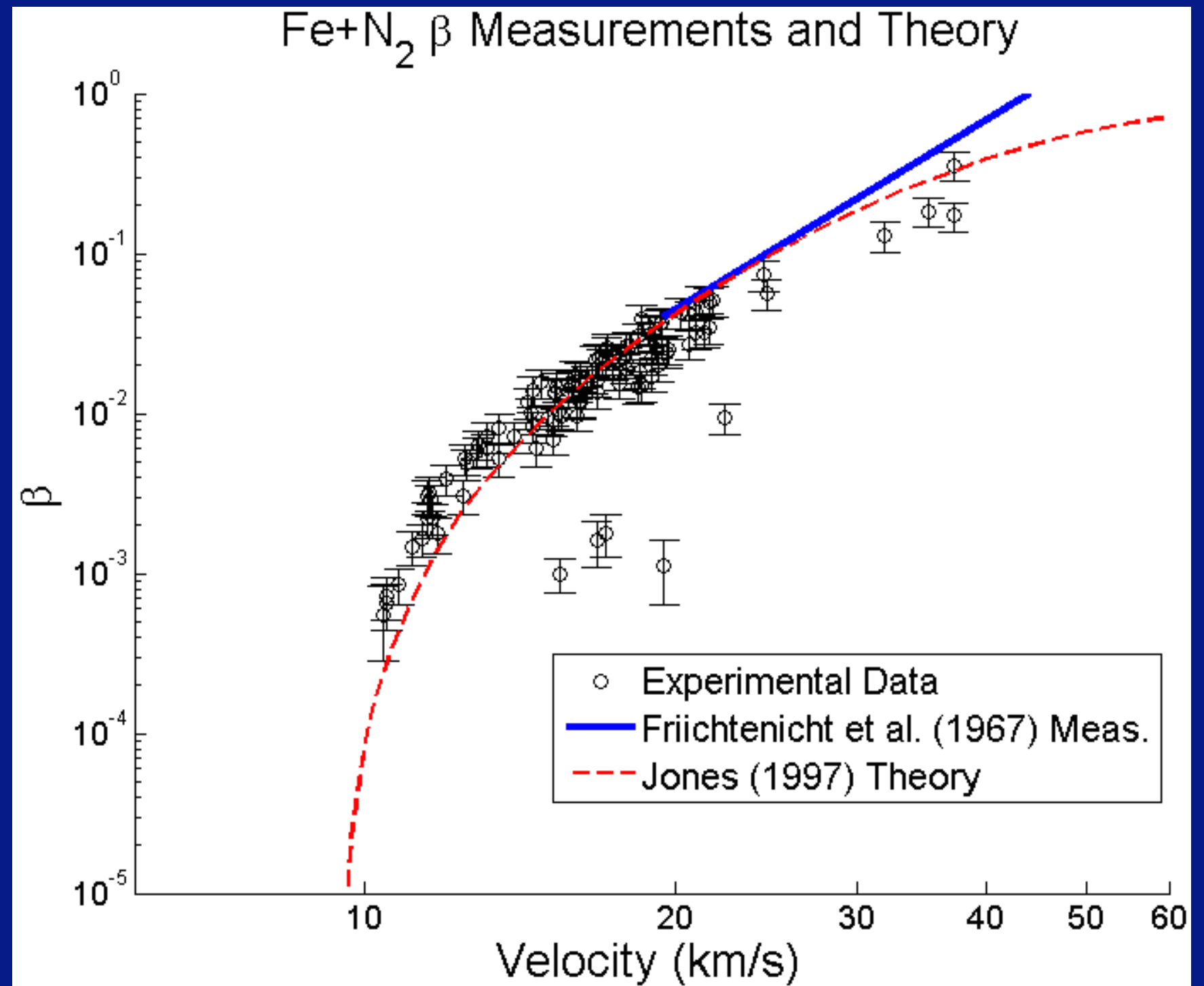
Charge Per Channel Shows Ablation Process



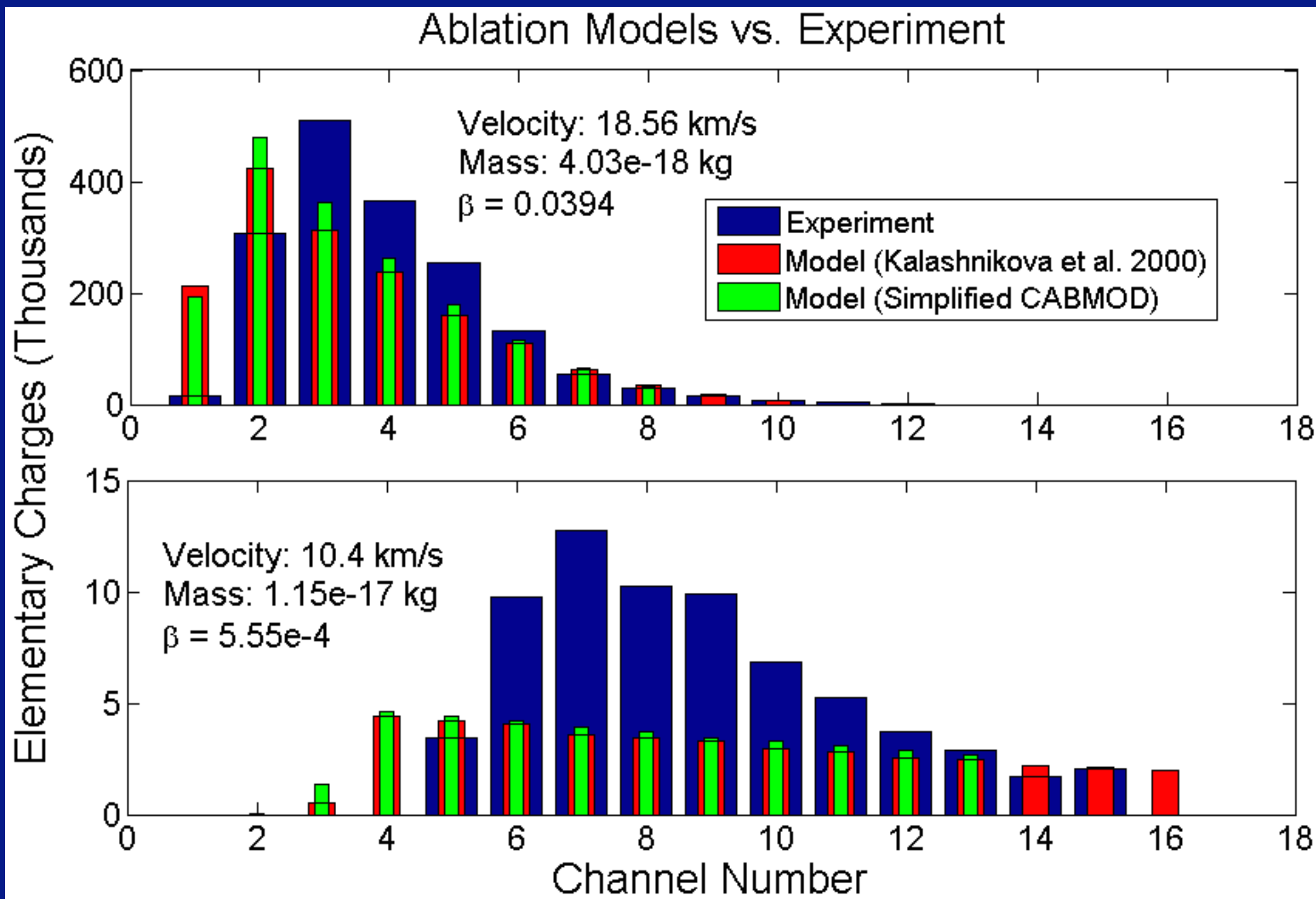
- The peak of each CSA signal gives the total charge for that channel.
- β is calculated by summing up all charge across 16 channels and dividing by the number of atoms in the particle.

First Result: β Measurements Extended to Lower Velocities

- The experimental $\beta(v)$ fits the Jones model for iron.
- First measurements below 20 km/s for Fe+N₂
- Data can also be used to verify ablation models

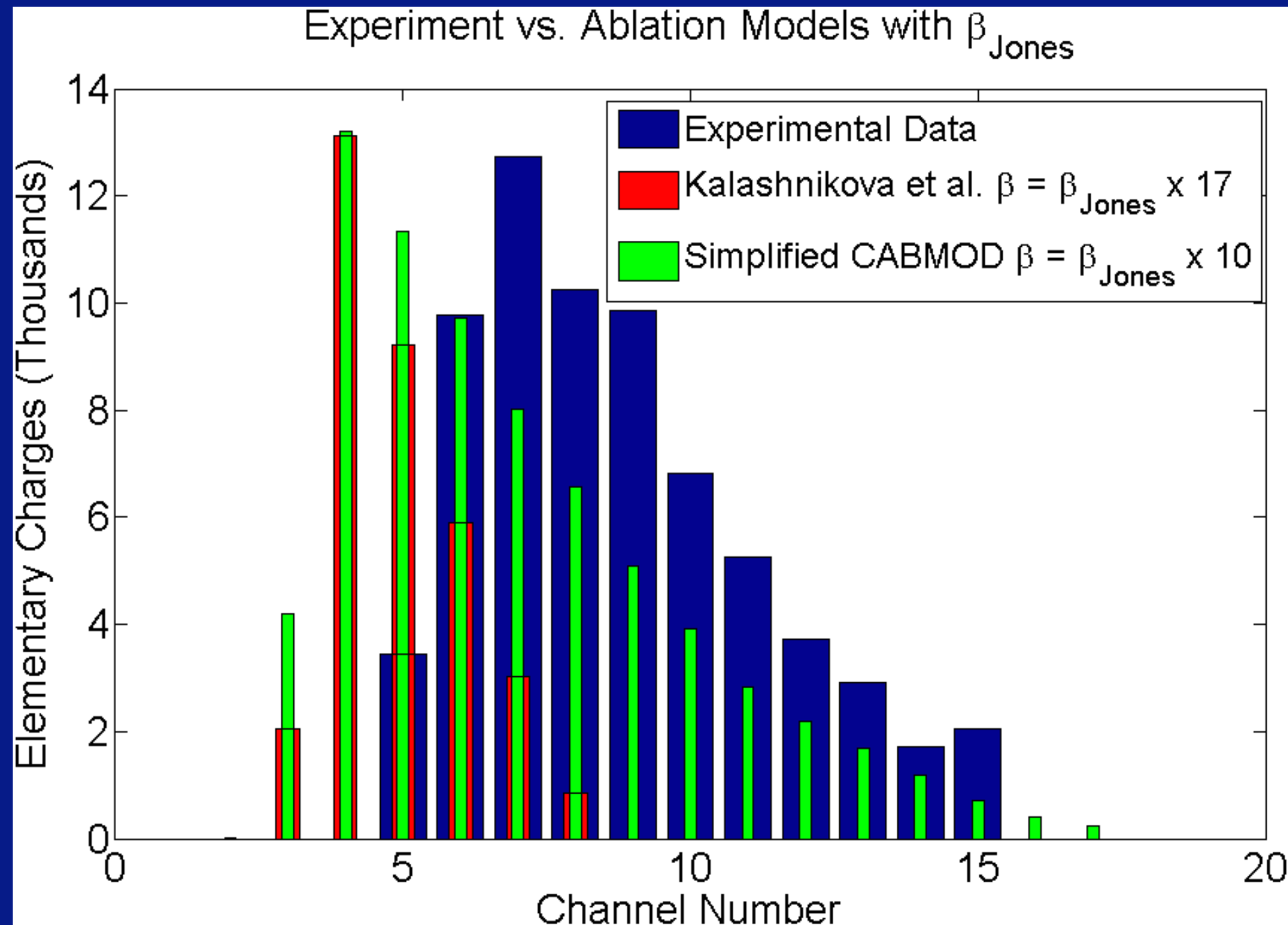


Compare Experiments to Ablation Models



Slow Speeds Present Unique Challenges

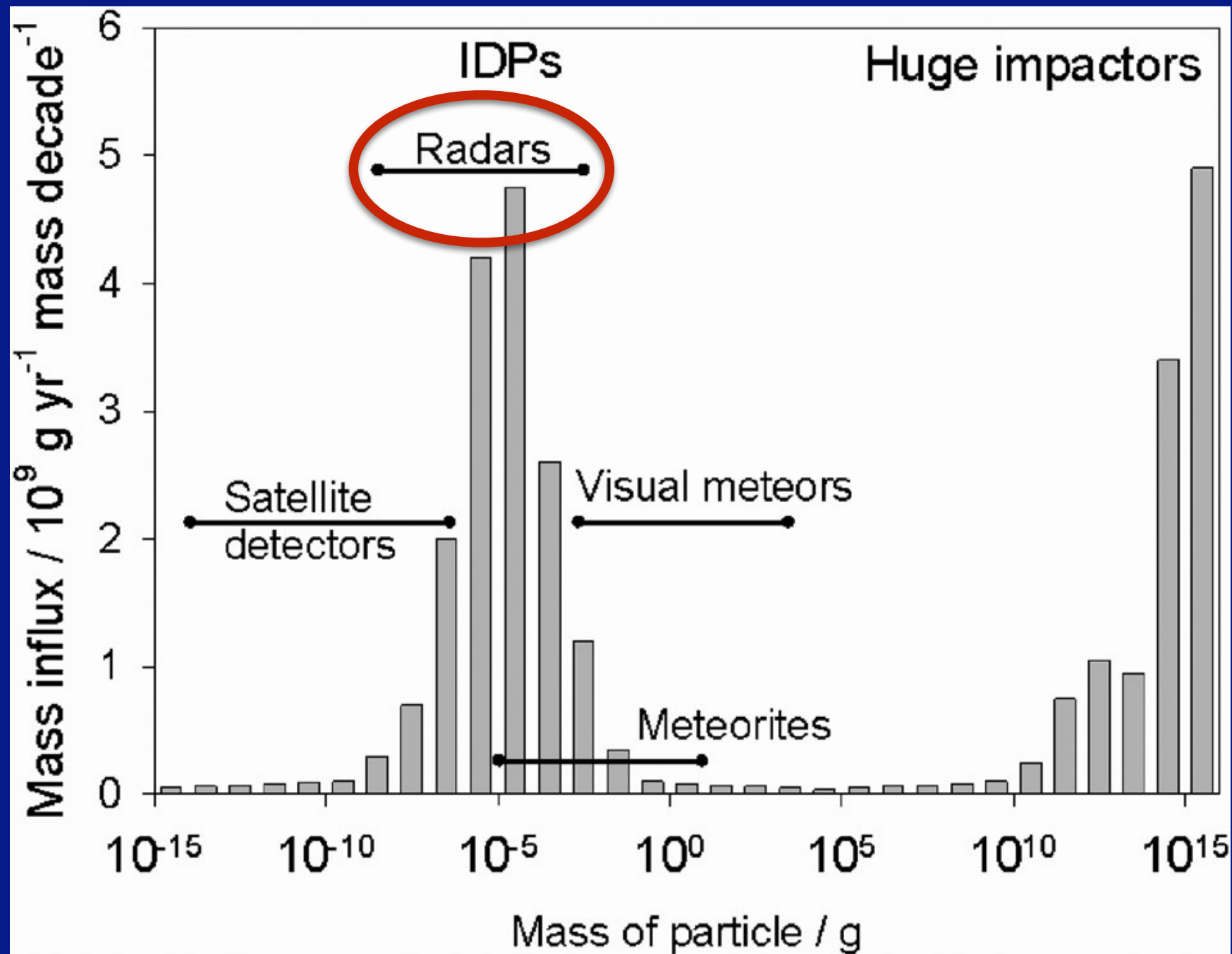
- Due to slowing of particle, used $\beta(v)$ from Jones (1997)
- Jones theory apparently requires adjusting at slow speeds.



Conclusions & Future Work

- Our results corroborate the Jones model for iron particles at high speeds and suggest it may need adjustment at slow speeds.
- We have already run experiments with different gases (O₂, air, He, CO₂) and those results are coming soon.
- We plan to use different meteor analogs (like olivine) instead of just using iron.
 - Potentially observe differential ablation.
- Future measurements are planned of the luminous efficiency.
 - Useful for visual meteors.

Meteor Radars Sensitive in Most Important Mass Range

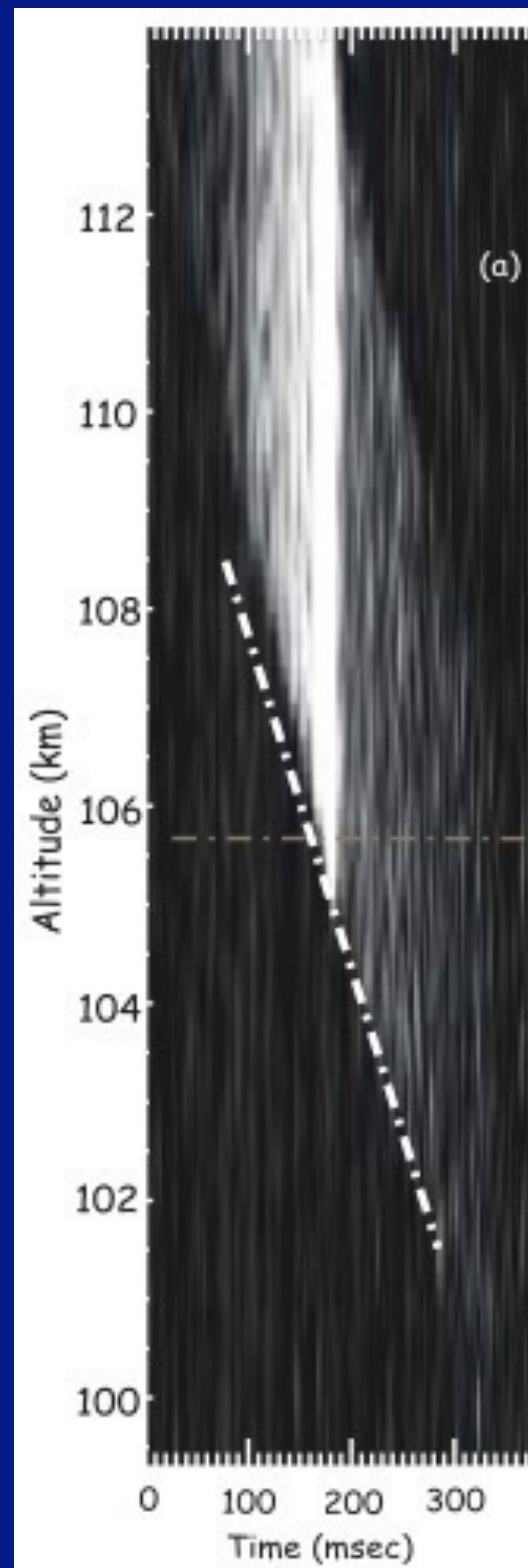


[Plane, 2012]

Radar Data Use Meteor Models

- Radar infers meteor characteristics from meteor plasma.

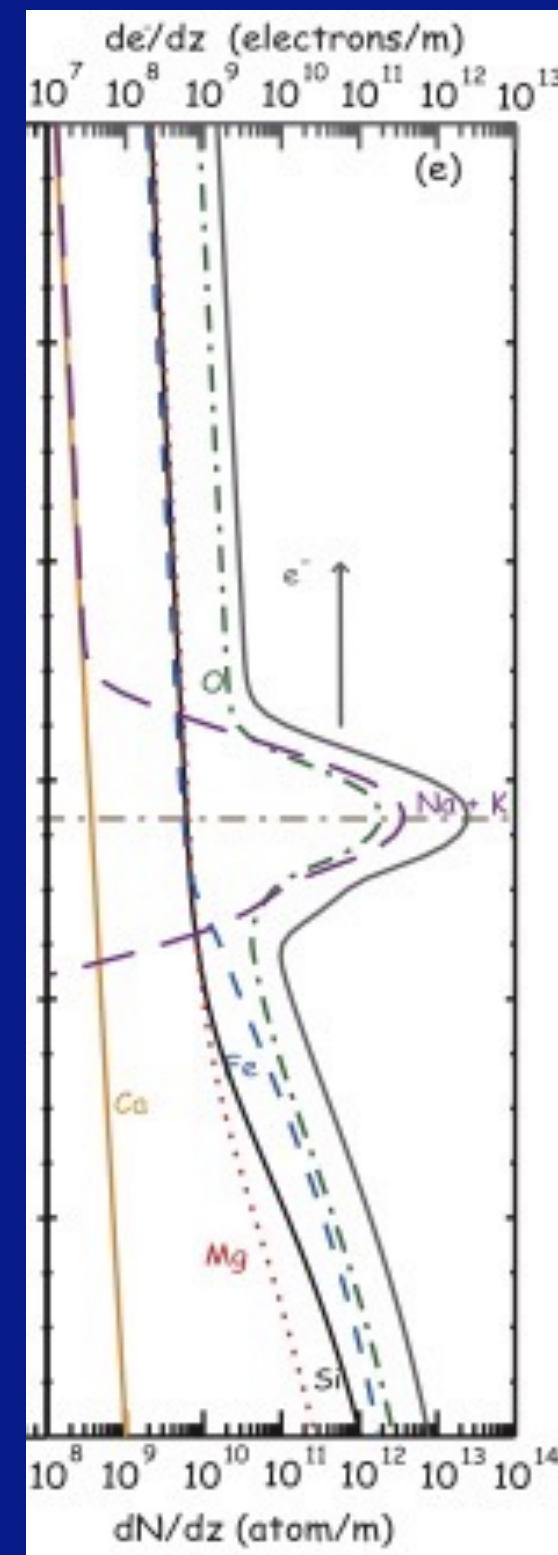
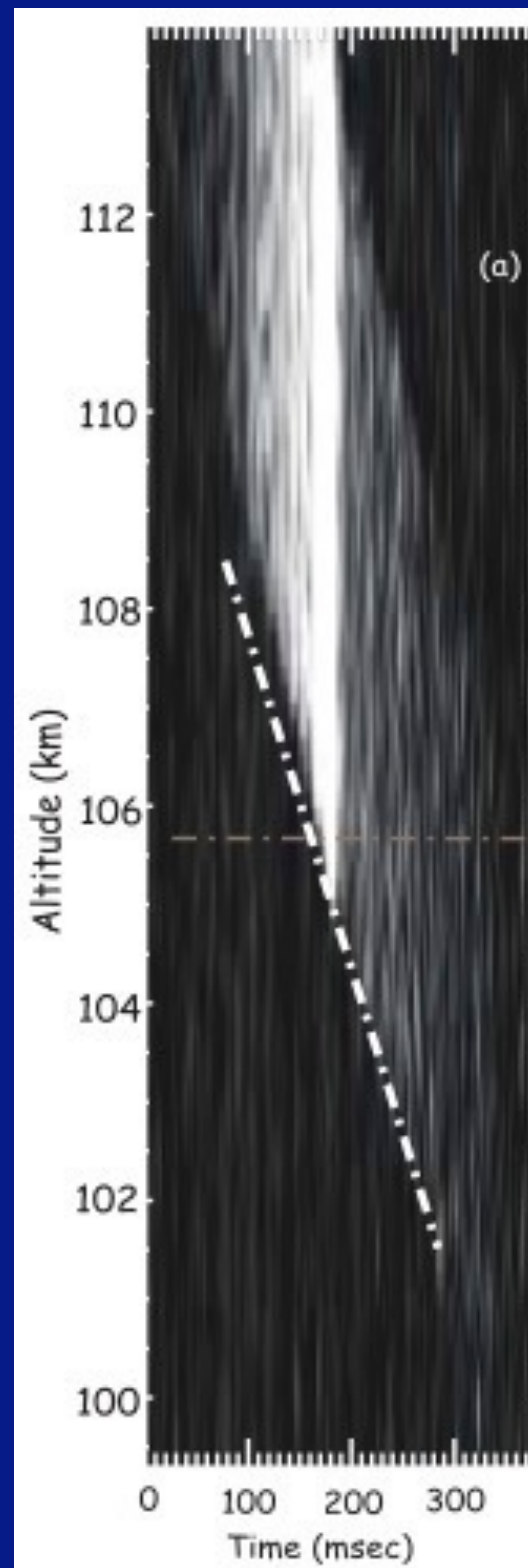
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CABMOD