



Consider a person whirling a ball on a string  
What force is being exerted?

**Obviously the person whirling the ball is exerting the force**

**This is called Centripetal force (center seeking)**

**The larger the mass, the larger the force**

$$F_c = \frac{mV^2}{r}$$

**The larger the velocity, the larger the force**

**The smaller the radius, the larger the force**

**For a planet rotating around the sun, the centripetal force pulling the planet in is the force of gravity. Thus, the gravitational force (given by Newton's law of universal gravitation) is equal to the centripetal force and for a planet of mass  $m$ :**

$$F_c = \frac{mV^2}{r} = \frac{mMG}{r^2}$$

**where  $r$  is the radius of the orbit and  $M$  is the mass of the sun**

**Can you get Kepler's 3rd law from the above?**

The velocity of the planet keeps it from crashing into the sun. We can easily calculate just the velocity necessary to keep the planet in orbit.

$$\frac{mV^2}{r} = \frac{mMG}{r^2}$$

$$V^2 = \frac{\cancel{m} MG \cancel{r}}{\cancel{r^2} \cancel{m}}$$

$$V^2 = \frac{MG}{r}$$

$$V = \sqrt{\frac{MG}{r}}$$

**With these equations, Newton was able to explain why Kepler's laws worked so well. His equations did not result from observation as did Kepler's. Instead they came from physical principles.**

**"Newton was the greatest genius who ever lived and the most fortunate for there cannot be more than once a world system to establish."**

**-J.J. Lagrange**

**The Giant Leap**

## **The Hubble Space Telescope**

**A Window on the Solar System and the Universe**

**Hubble's Greatest hits**

**Comet Shoemaker-Levey Bombards Jupiter**

**Storm on Saturn's Equator**

**Martian Weather-Ice Caps and Clouds**

# **The Galileo Mission to Jupiter**

**Pictures from galileo**

**Jupiter's Moon Io**

**Jupiter's Moon Io -2**

# Continental Drift

The following cards show continental drift.

The first card shows the arrangements of the continents as they are today.

The start button takes you backward in time to the super continent of Pangaea 200 million years ago.

750 million years ago

Snapshots



# Plate Tectonics

**Map of Earth's Plates**

**Earthquake Frequency**

**California Faults**

**San Andreas Fault**

# What is the Fate of the Solar System?

Life of Earth

Red Giant

Sun is born

Life on Earth

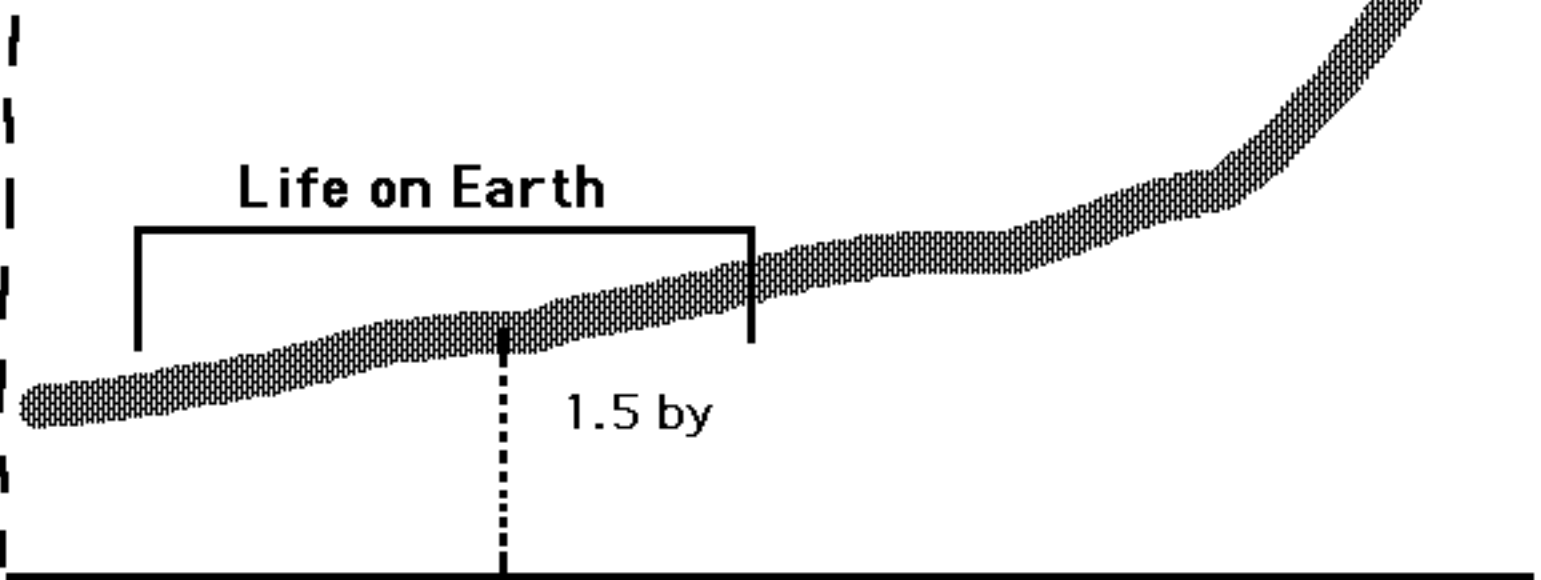
1.5 by

-4.5

0

+7.5

Billions of Years



**Problem 9 Page 27 Calculate the velocity necessary to keep the moon in a stable orbit around the Earth at a distance of 383,000 Km**

**Use the formula  $V^2 = MG/r$**

**M=Mass of Earth ( $5.97 \times 10^{24}$  Kg)**

**r must be in meters  $= (3.83 \times 10^5 \text{ Km})(10^3 \text{ m/Km}) = 3.83 \times 10^8 \text{ m}$**

**$V^2 = 6.67 \times 10^{-11} \text{ Nm}^2/\text{Kg}^2) (5.97 \times 10^{24} \text{ Kg}) / 3.83 \times 10^8 \text{ m}$**

**$V^2 = 1.04 \times 10^6 \text{ Nm/Kg}$**

**Units are a bit tricky here  $1 \text{ N} = 1 \text{ Kg m/sec}^2$**

**$V^2 = 1.04 \times 10^6 \text{ Kg m}^2/\text{Kgsec}^2 = 1.04 \times 10^6 \text{ m}^2/\text{sec}^2$**

**Now take the square root of  $1.04 \times 10^6 \text{ m}^2/\text{sec}^2$**

**$V = 1.02 \times 10^3 \text{ m/sec}$**